

# A String-Theoretic Introduction to Mirror Symmetry

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## Warnings

These slides serve as a motivational introduction to mirror symmetry from a physical perspective. We will go through some basic concepts aimed at mathematical audiences. The logic flow is more chronological than pedagogical.

I am new to this subject, so mistakes and missings are inevitable.

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### Definition (Action Functional)

Let  $\mathcal{M}$  be the “space of configurations” of a physical system. The action functional is a linear map  $S: \mathcal{M} \rightarrow \mathbb{R}$ .

In classical mechanics,  $\mathcal{M}$  is the space of piecewise smooth paths from  $x$  to  $y$  on a smooth manifold  $M$ .

In classical field theory,  $\mathcal{M}$  is the space of sections of a vector bundle  $E$  over the spacetime manifold  $(M, g)$ .

We need a well-behaved measure on  $\mathcal{M}$  to perform calculus of variations and integrations. This is usually not well-defined, but physicists just assume it.

### Proposition (Principle of Least Action)

*The trajectory of a system is the one which extremise the action:  
 $\delta S = 0$ .*

In classical mechanics, the action arises as the integral of the Lagrangian:

$$S = \int_{t_0}^{t_1} L(q_a, \dot{q}_a, t) dt.$$

The principle of least action is equivalent to the Euler–Lagrange equations (“equations of motions”):

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{q}_a} - \frac{\partial L}{\partial q_a} = 0,$$

where  $(q_1, \dots, q_n)$  is a set of coordinates on  $M$ .

### Proposition (Principle of Least Action)

*The trajectory of a system is the one which extremise the action:  
 $\delta S = 0$ .*

In classical field theory, the action arises as the integral of the Lagrangian density:

$$S = \int_M \mathcal{L}(\varphi_\alpha, \partial_a \varphi_\alpha) \, d \operatorname{vol}_g = \int_{\mathbb{R}^{1,d-1}} \mathcal{L} \sqrt{-\det g} \, d^n x.$$

The Euler–Lagrange equations:

$$\sum_a \frac{\partial}{\partial x^a} \frac{\partial \mathcal{L}}{\partial (\partial_a \varphi_\alpha)} - \frac{\partial \mathcal{L}}{\partial \varphi_\alpha} = 0.$$

## Path Integral Formalism

Now we move from classical to quantum! In the path integral formalism, the primary object is the partition function:

$$Z := \int_{\mathcal{M}} \mathcal{D}[\gamma] e^{iS(\gamma)/\hbar},$$

where  $\mathcal{D}[\gamma]$  is a well-behaved measure on  $\mathcal{M}$  (which, unfortunately, does not exist in general. This is one of the most notable mathematical difficulties of quantum field theory.)

In the classical limit  $\hbar \rightarrow 0$ , only the classical solution such that  $\delta S = 0$  contributes to the partition function.

An observable  $O$  is an operator-valued distribution on  $\mathcal{M}$  with the expectation:

$$\langle O \rangle := \int_{\mathcal{M}} \mathcal{D}[\gamma] O(\gamma) e^{iS(\gamma)/\hbar}.$$



## Polyakov Action

A string is a 1-dimensional object in the space. It traces out a 2-dimensional surface (the “worldsheet”) in the spacetime  $M$ . So mathematically, string theory is about (the quantisation of) the embedding  $X: \Sigma \rightarrow M$  of a Lorentzian surface into a  $d$ -dimensional Lorentzian manifold.

### Definition (Polyakov Action)

The classical bosonic string is described by the Polyakov action:

$$S[X, h] = -\frac{1}{4\pi\alpha'} \int_{\Sigma} d^2\sigma \sqrt{-\det h} h^{\alpha\beta} \partial_{\alpha} X^{\mu} \partial_{\beta} X^{\nu} g_{\mu\nu},$$

where  $h, g$  are the metrics on  $\Sigma$  and  $M$  respectively, and  $X = (X^{\mu})$  are the coordinates of  $M$ .

## Supersymmetry

Slogan:

bosonic fields	fermionic fields
even variables ("c-numbers")	odd variables ("Grassmann numbers")
commutator $[a, b] = ab - ba$	anti-commutator $\{a, b\} = ab + ba$

ungraded Lie algebra  $\rightarrow \mathbb{Z}/2$ -graded Lie algebra  
("supersymmetric algebra")

Supersymmetric scalar multiplet: (scalar fields  $X^\mu$ , fermions  $\psi^\mu$ );

Supergravity multiplet: (frame fields  $e_\alpha^a$ , gravitini  $\chi_\alpha$ ).

## Ramond–Neveu–Schwarz Strings

Bosonic string +  $N = 1$  worldsheet supersymmetry = RNS string.

### Proposition (RNS String Action)

*The RNS string (fixed to superconformal gauge) is described by the action*

$$S = -\frac{1}{8\pi} \int_{\Sigma} d^2\sigma \left( \frac{2}{\alpha'} \partial_{\alpha} X^{\mu} \partial^{\alpha} X^{\nu} + 2i\bar{\psi}^{\mu} \rho^{\alpha} \partial_{\alpha} \psi^{\nu} \right) g_{\mu\nu}$$

The action is invariant under the supersymmetric transformation:

$$\begin{aligned} \sqrt{\frac{2}{\alpha'}} \delta_{\epsilon} X^{\mu} &= i\bar{\epsilon} \psi^{\mu} \\ \delta_{\epsilon} \psi^{\mu} &= \sqrt{\frac{2}{\alpha'}} \frac{1}{2} \rho^{\alpha} \partial_{\alpha} X^{\mu} \epsilon \end{aligned}$$

## Conformal Field Theory

The classical string action is Weyl invariant:

$$h \mapsto \Omega^2 h$$

So string theory is a 2-dimensional conformal field theory!

The stress-energy tensor is traceless:  $\text{tr } T = 0$ .

After quantisation, Weyl invariance must be preserved:

$$\langle \text{tr } T \rangle = -\frac{c}{12} R \quad (\text{worldsheet scalar curvature})$$

So *central charge*  $c = 0$ .

After massaging operator product expansions...

$$c = d - 10 \implies d = 10$$

Conformal field theory tells us that superstrings live in 10 spacetime dimensions!!! Needs compactification of 6 extra dimensions to get real-world physics. See later.

## Conformal Field Theory

For closed strings, the periodic boundary condition on  $\psi^\mu$ :

$$\psi^\mu(\sigma) = +\psi^\mu(\sigma + \ell), \quad \text{Ramond sector}$$

$$\psi^\mu(\sigma) = -\psi^\mu(\sigma + \ell), \quad \text{Neveu-Schwarz sector}$$

Two independent spinors  $\psi_+^\mu(\sigma^+)$ ,  $\psi_-^\mu(\sigma^-)$ , so 4 types of closed strings: (R,R), (NS,NS), (R,NS), (NS,R).

$N = 1$  super-Virasoro algebra:

$$[L_m, L_n] = (m - n)L_{m+n} + \frac{c}{12}m(m^2 - 1)\delta_{m,-n}$$

$$[L_m, G_r] = \left(\frac{m}{2} - r\right)G_{m+r}$$

$$\{G_r, G_s\} = 2L_{r+s} + \frac{c}{12}(4r^2 - 1)\delta_{r,-s}$$

Ramond sector:  $r, s \in \mathbb{Z}$ ; Neveu-Schwarz sector:  $r, s \in \mathbb{Z} + \frac{1}{2}$ .

## Working towards Superstring Theories

Recall that in bosonic string theory, critical dimension  $d = 26$ ; and the ground state is *tachyonic* ( $m^2 < 0$ ).

For superstring with worldsheet as a Riemann surface:

Modular invariance & Vanishing of one-loop partition function



GSO projection  $(-1)^F, (-1)^{\bar{F}}$



Tachyon-free & Spacetime supersymmetry  $\mathcal{N} = 2$

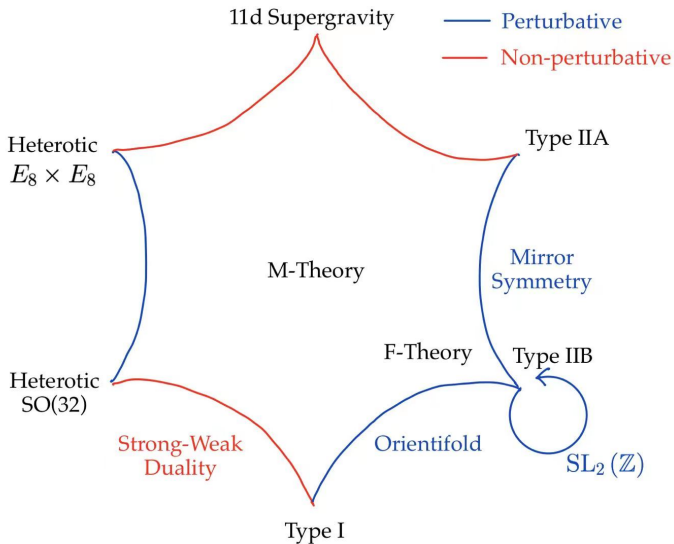


Type IIA superstring theory  
(Non-chiral)



Type IIB superstring theory  
(Chiral)

# The Landscape of String Dualities



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## Compactification

Superstring theory works in 10 dimensions, but our physical world has only 4 dimensions. We would like to factor the spacetime as

$$\mathbb{R}^{1,3} \times M_6$$

where  $M_6$  is a 6-dimensional compact Riemannian manifold of length scale  $\sim 10^{-35}$  m.

Even though we do not observe  $M_6$ , the geometry of it actually determines the physics in the Minkowski space  $\mathbb{R}^{1,3}$ !

Toroidal compactification:  $M_6 = T^6$ .

All supersymmetries preserved  $\rightarrow$  phenomenologically unfavourable.

Calabi–Yau compactification of heterotic strings:  $M_6 = \text{CY}_3$ .

$\mathcal{N} = 1$  minimal supersymmetric standard model (MSSM)!

There are simply too many choices of  $M_6$  for string theory to make physical predictions. This is why string theory is not considered as a part of physics by many people.

## Naive Example: Toroidal Compactification

Consider the compactification  $M^{d+1} = M^d \times S^1$ :

$$x^d \sim x^d + 2\pi R w.$$

$R \in \mathbb{R}$  is the radius of the circle  $S^1$ ;  $w \in \mathbb{Z}$  is the winding number.

Slogan: Gravity in  $(d+1)$ -dimensions produces electromagnetism in  $d$ -dimensions.

Factorisation of metric:

$$G = \sum_{M,N=0}^d g_{MN} dx^M dx^N = \sum_{\mu,\nu=0}^{d-1} g_{\mu\nu} dx^\mu dx^\nu + g_{dd} \left( dx^d + \sum_{\mu=0}^{d-1} A_\mu dx^\mu \right)^2.$$

Coordinate change  $x^d \mapsto x^d + \lambda(x^\mu)$  produces gauge transformation  $A_\mu \mapsto A_\mu - \partial_\mu \lambda$ .

“**Kaluza–Klein Reduction**”.  $A_\mu$ : KK gauge boson.

## Naive Example: Toroidal Compactification

Scalar field  $\phi$  in  $(d + 1)$ -dimensions has the mode expansion:

$$\phi(x^M) = \sum_{n \in \mathbb{Z}} \phi_n(x^\mu) e^{inx^d/R}.$$

The momentum  $p_d = n/R$  is quantised.

Mass for bosonic string:

$$m^2 = \frac{n^2}{R^2} + \frac{w^2 R^2}{\alpha'^2} + \frac{2}{\alpha'} (N + \tilde{N} - 2), \quad N - \tilde{N} = nw.$$

T-duality:

$$R \leftrightarrow \frac{\alpha'}{R}, \quad n \leftrightarrow w.$$

Also Type IIA  $\leftrightarrow$  Type IIB.

## Naive Example: Toroidal Compactification

Now consider compactification  $M^{d+D} = M^d \times T^D$ ,  $T^D := \mathbb{R}^D / 2\pi\Lambda_D$ .  
*For modular invariance the lattice  $\Lambda_D$  must be self-dual:  $\Lambda_D = \Lambda_D^*$ .*

KK reduction  $\implies$  Gauge boson  $A_\mu$ , 2-form  $B_{mn}$ .

$B$ -field: “massless scalar degrees of freedom contributed by the internal 2-form” ...

T-duality:

$$\frac{1}{\alpha'}(\mathbf{g} + \mathbf{b}) \leftrightarrow \alpha'(\mathbf{g} + \mathbf{b})^{-1}, \quad \mathbf{n} \leftrightarrow \mathbf{w}$$

Moduli space:

$$\frac{\mathrm{O}(D, D)}{\mathrm{O}(D) \times \mathrm{O}(D)} \Big/ \mathrm{O}(D, D; \mathbb{Z}).$$

## Naive Example: Toroidal Compactification

2-torus:  $T^2 = \mathbb{R}^2 / (\mathbb{Z}e_1 + \mathbb{Z}e_2)$ .  $g_{ij} := e_i \cdot e_j$ .

Complexified Kähler modulus:  $T := \frac{1}{\alpha'} (B + i\sqrt{\det g}) = T_1 + iT_2$ .

Complex structure modulus:

$$U := \frac{\|e_2\|}{\|e_1\|} e^{i\varphi(e_1, e_2)} = \frac{g_{12} + i\sqrt{\det g}}{g_{11}} = U_1 + iU_2.$$

$$(g_{ij}) = \alpha' \frac{T_2}{U_2} \begin{pmatrix} 1 & U_1 \\ U_1 & |U|^2 \end{pmatrix}.$$

T-duality:  $(n_1, n_2, w_1, w_2) \leftrightarrow (-w_1, n_2, -n_1, w_2)$ ;  $T \leftrightarrow U$ .

This is a prototype of mirror symmetry, where the complex structure and the Kähler structure interchanges under the mirror transformation.

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Compactification:  $\mathbb{R}^{1,3} \times M_6$ . Why Calabi–Yau?

Ricci-flat:  $R_{mn} = 0$ . (*Think of vacuum solutions of Einstein equation.*)

SU(3) holonomy.

### Definition (Curvature)

Let  $(M, g)$  be a Riemannian/Lorentzian manifold with Levi-Civita connection  $\nabla$ . The Riemann curvature is given by  $R_{abc}{}^d \partial_d = [\nabla_a, \nabla_b] \partial_c$ . The Ricci curvature is given by  $R_{ab} := R_{cab}{}^c$ .

### Definition (Holonomy Group)

Let  $(M, g)$  as above. The holonomy group  $\text{Hol}_x(\nabla)$  based at  $x \in M$  is the group generated by the parallel transports  $P_\gamma$ , where  $\gamma$  is a loop in  $M$  based at  $x$ .

Compactification:  $\mathbb{R}^{1,3} \times M_6$ . Why Calabi–Yau?

Ricci-flat:  $R_{mn} = 0$ .

SU(3) holonomy.

Decomposition of Weyl representation under  $\text{SO}(1, 9) \rightarrow \text{SO}(1, 3) \times \text{SO}(6)$ :

$$\mathbf{16} = \mathbf{2}_L \otimes \bar{\mathbf{4}} \oplus \mathbf{2}_R \otimes \mathbf{4}.$$

Further decomposition of  $\mathbf{4}$  of  $\mathfrak{so}(6) \cong \mathfrak{su}(4)$  if  $M_6$  has a SU(3)-structure:

$$\mathbf{4}_{\mathfrak{su}(4)} = (\mathbf{1} \oplus \mathbf{3})_{\mathfrak{su}(3)}.$$

The singlet state of  $\mathfrak{su}(3)$  is a covariantly constant spinor of  $M_6$ , i.e.  $\nabla_m \epsilon = 0$ . This produces  $\mathcal{N} = 1$  supersymmetry in  $d = 4$ .

Side effect:  $M_6$  is Ricci-flat. (Exercise: try to prove this~)



## Kähler Manifolds

Slogan: A Kähler manifold is the one with compatible Riemannian, complex, and symplectic structures, encoded by:

$$\omega(X, Y) = g(JX, Y),$$

where:

$\omega \in \Gamma(\wedge^2 T^*M)$  is a symplectic form;

$J \in \Gamma(\text{End}(TM))$  is a complex structure ( $J^2 = -\text{id}$ );

$g \in \Gamma(S^2 T^*M)$  is a Riemannian metric.

Any two of  $(\omega, J, g)$  determines the third one.

Local expression:

$$\omega = ig_{\mu\bar{\nu}} dz^\mu \wedge d\bar{z}^\nu, \quad g_{\mu\nu} = g_{\bar{\mu}\bar{\nu}} = 0.$$

## Hodge Theory

Kähler manifolds are a special class of complex manifolds (real manifolds with an integrable almost complex structure  $J$ .)

Exterior differential splits into holomorphic & anti-holomorphic parts:

$$d = \partial + \bar{\partial}.$$

de Rham cohomology  $H_{\text{dR}}^n(M; \mathbb{C}) = \bigoplus_{p+q=n} H^{p,q}(M)$

Dolbeault cohomology  $H^{p,q}(M) = H^q(M, \Omega_M^p)$ .

$$\partial: H^{p,q}(M) \rightarrow H^{p+1,q}(M); \quad \bar{\partial}: H^{p,q}(M) \rightarrow H^{p,q+1}(M).$$

Hodge numbers:  $h^{p,q}(M) := \dim_{\mathbb{C}} H^{p,q}(M)$ .

Serre duality:

$$H^{p,q}(M) \cong H^{n-q,n-p}(M).$$

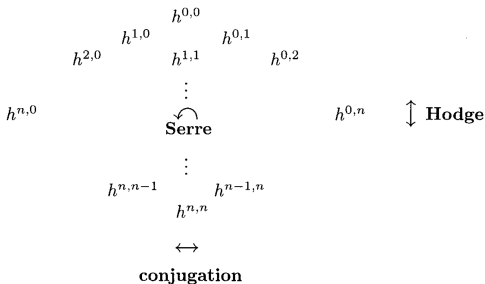
Hodge star:

$$H^{p,q}(M) \cong H^{n-p,n-q}(M)$$

Complex conjugation:

$$H^{p,q}(M) \cong H^{q,p}(M)$$

Hodge diamond:



There are a lot of different and inequivalent definitions of a Calabi–Yau manifold.

### Definition (Calabi–Yau Manifolds)

Let  $M$  be a  $2n$ -dimensional compact Kähler manifold. We say that  $M$  is Calabi–Yau, if it satisfies any of the following conditions:

1.  $M$  has vanishing first Chern class  $c_1(M)$ ;
2.  $M$  admits a Ricci-flat Kähler metric;
3.  $M$  has trivial canonical bundle  $K_M = \bigwedge^n \Omega_M$ ;
4.  $M$  admits a Kähler metric with holonomy contained in  $SU(n)$ .
5.  $M$  has a unique nowhere vanishing holomorphic  $n$ -form.

The conditions satisfy (1)  $\iff$  (2)  $\iff$  (3)  $\iff$  (4)  $\iff$  (5).

They are all equivalent if  $M$  is simply connected.

The hard part is (1)  $\implies$  (2), known as the *Calabi conjecture*, proven by Yau in 1978.

## Definition (Physicists' Calabi–Yau Manifolds)

Let  $M$  be a  $2n$ -dimensional compact Kähler manifold. We say that  $M$  is Calabi–Yau, if  $M$  admits a Kähler metric with holonomy *exactly equal to*  $SU(n)$ .

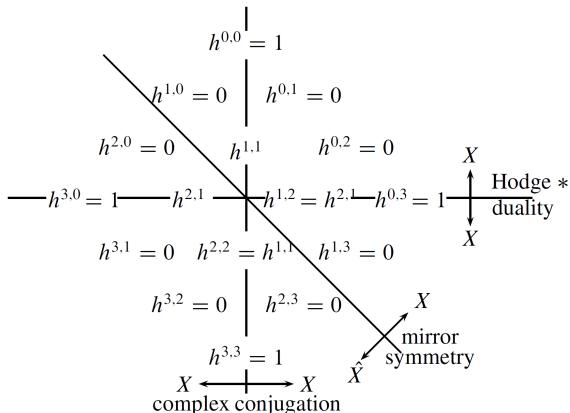
Hodge numbers:

$$h^{0,0} = 1, \quad h^{n,0} = 1, \quad h^{i,0} = 0 \text{ for } 0 < i < n.$$

So for a Calabi–Yau 3-fold, the only independent Hodge numbers are  $h^{1,1}$  and  $h^{1,2}$ .

Euler characteristic  $\chi(\text{CY}_3) = 2(h^{1,1}(\text{CY}_3) - h^{1,2}(\text{CY}_3))$ .

## Mirror Calabi–Yau 3-folds



Mirror pair  $(X, X^\vee)$ :

$$h^{1,1}(X) = h^{1,2}(X^\vee), \quad h^{1,2}(X) = h^{1,1}(X^\vee).$$

Infinitesimal deformation of the metric  $g \mapsto g + \delta g$  preserving Ricci-flatness.

Lichnerowicz equation:

$$\Delta_L \delta g_{\mu\nu} := \nabla^\rho \nabla_\rho \delta g_{\mu\nu} + 2R_{\mu}{}^\rho{}_\nu{}^\sigma \delta g_{\rho\sigma} = 0$$

1. Mixed indices  $\delta g_{i\bar{j}}$ :

$$(\Delta \delta g)_{i\bar{j}} = 0.$$

So  $\delta g_{i\bar{j}}$  are components of a  $(1, 1)$ -form.

$$\delta g_{i\bar{j}} = \sum_{\alpha=1}^{h^{1,1}} \tilde{t}^\alpha b_{i\bar{j}}^\alpha$$

where  $\tilde{t}^\alpha \in \mathbb{R}$  are the Kähler moduli, which spans the *Kähler cone*.

Infinitesimal deformation of the metric  $g \mapsto g + \delta g$  preserving Ricci-flatness.

Lichnerowicz equation:

$$\Delta_L \delta g_{\mu\nu} := \nabla^\rho \nabla_\rho \delta g_{\mu\nu} + 2R_{\mu}{}^\rho{}_\nu{}^\sigma \delta g_{\rho\sigma} = 0$$

1. Combining with the deformation of  $B$ -field:

$$(i\delta g_{i\bar{j}} + \delta B_{i\bar{j}}) dz^i \wedge d\bar{z}^j = \sum_{\alpha=1}^{h^{1,1}} t^\alpha b_{i\bar{j}}^\alpha$$

Now  $t^\alpha \in \mathbb{C}$  are complexified Kähler moduli.



Infinitesimal deformation of the metric  $g \mapsto g + \delta g$  preserving Ricci-flatness.

Lichnerowicz equation:

$$\Delta_L \delta g_{\mu\nu} := \nabla^\rho \nabla_\rho \delta g_{\mu\nu} + 2R_{\mu}{}^\rho{}_\nu{}^\sigma \delta g_{\rho\sigma} = 0$$

2. Pure indices  $\delta g_{i\bar{j}}$ :

$$\Delta_{\bar{\partial}} \delta g^i = 0, \quad \delta g^i := g^{i\bar{k}} \delta g_{\bar{k}\bar{j}} d\bar{z}^{\bar{j}}$$

$\delta g^i$  is a  $\mathbb{T}_M := \mathbb{T}^{1,0} M$ -valued  $(0, 1)$ -form.

$H_{\bar{\partial}}^{0,1}(M, \mathbb{T}_M) \cong H^1(M, \mathbb{T}_M) \cong H^{2,1}(M)$ , so

$$\Omega_{ijk} \delta g_{\bar{l}}^k = \sum_{a=1}^{h^{2,1}} t^a b_{ij\bar{l}}^\alpha$$

where  $t^\alpha \in \mathbb{C}$  are the complex structure moduli.

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## Massless Spectra of Type IIA & Type IIB Theories

Type IIA on  $X \longleftrightarrow$  Type IIB on  $X^\vee!$

**Table 14.1** Massless spectrum of the type II theories on  $CY_3$

Multiplet	Component fields	Multiplicity
<b>Type IIA</b>		
Gravity	$g_{\mu\nu}, \psi_{\mu\alpha}\eta, \tilde{\psi}_{\mu\dot{\alpha}}\eta, \bar{\psi}_{\mu\dot{\alpha}}\eta_{\bar{i}\bar{j}\bar{k}}, \tilde{\psi}_{\mu\alpha}\eta_{\bar{i}\bar{j}\bar{k}}, (C_1)_\mu$	1
Hyper	$\lambda_\alpha\eta, \tilde{\lambda}_{\dot{\alpha}}\eta, \bar{\lambda}_{\dot{\alpha}}\eta_{\bar{i}\bar{j}\bar{k}}, \tilde{\lambda}_\alpha\eta_{\bar{i}\bar{j}\bar{k}}, \Phi, B_{\mu\nu}, (C_3)_{ijk}, (C_3)_{\bar{i}\bar{j}\bar{k}}$	1
Hyper	$\psi_\alpha\eta_{i,\bar{j}\bar{k}}, \bar{\psi}_{\dot{\alpha}}\eta_{\bar{i},\bar{j}}, \tilde{\psi}_{\dot{\alpha}}\eta_{i,\bar{j}\bar{k}}, \tilde{\psi}_\alpha\eta_{\bar{i},\bar{j}}, g_{ij}, g_{\bar{i}\bar{j}}, (C_3)_{i\bar{j}\bar{k}}, (C_3)_{\bar{i}jk}$	$h^{2,1}$
Vector	$(C_3)_{\mu i\bar{j}}, \bar{\psi}_{\dot{\alpha}}\eta_{i,\bar{j}}, \psi_\alpha\eta_{\bar{i},\bar{j}\bar{k}}, \tilde{\psi}_\alpha\eta_{i,\bar{j}}, \tilde{\psi}_{\dot{\alpha}}\eta_{\bar{i},\bar{j}\bar{k}}, g_{i\bar{j}}, B_{i\bar{j}}$	$h^{1,1}$
<b>Type IIB</b>		
Gravity	$g_{\mu\nu}, \bar{\psi}_{\mu\dot{\alpha}}\eta, \psi_{\mu\alpha}\eta_{\bar{i}\bar{j}\bar{k}}, \tilde{\psi}_{\mu\dot{\alpha}}\eta, \tilde{\psi}_{\mu\alpha}\eta_{\bar{i}\bar{j}\bar{k}}, (C_4^+)_{\mu ijk}$	1
Hyper	$\lambda_\alpha\eta, \bar{\lambda}_{\dot{\alpha}}\eta_{\bar{i}\bar{j}\bar{k}}, \tilde{\lambda}_\alpha\eta, \tilde{\lambda}_{\dot{\alpha}}\eta_{\bar{i}\bar{j}\bar{k}}, \Phi, a, B_{\mu\nu}, (C_2)_{\mu\nu}$	1
Hyper	$\psi_\alpha\eta_{i,\bar{j}}, \bar{\psi}_{\dot{\alpha}}\eta_{\bar{i},\bar{j}\bar{k}}, \tilde{\psi}_\alpha\eta_{i,\bar{j}}, \tilde{\psi}_{\dot{\alpha}}\eta_{\bar{i},\bar{j}\bar{k}}, g_{i\bar{j}}, B_{i\bar{j}}, (C_2)_{i\bar{j}}, (C_4^+)_{\mu vi\bar{j}}$	$h^{1,1}$
Vector	$(C_4^+)_{\mu i\bar{j}\bar{k}}, \bar{\psi}_{\dot{\alpha}}\eta_{i,\bar{j}\bar{k}}, \psi_\alpha\eta_{\bar{i},\bar{j}}, \tilde{\psi}_{\dot{\alpha}}\eta_{i,\bar{j}\bar{k}}, \tilde{\psi}_\alpha\eta_{\bar{i},\bar{j}}, g_{ij}, g_{\bar{i}\bar{j}}$	$h^{2,1}$